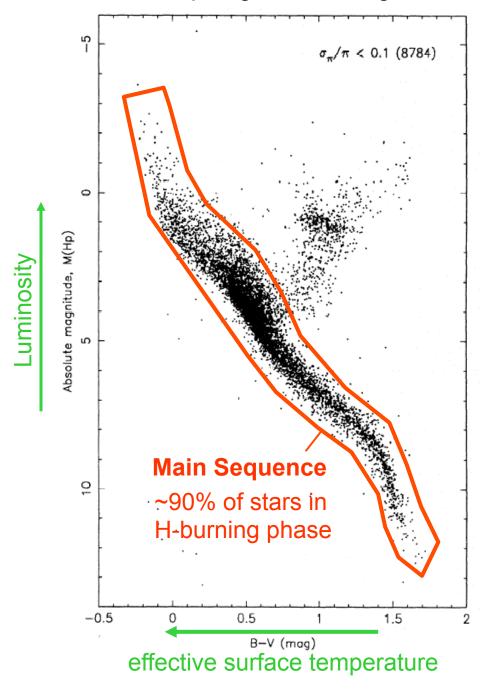
Hertzsprung-Russell diagram



Perryman et al. A&A 304 (1995) 69 HIPPARCOS distance measurements

Magnitude:

Measure of stars brightness

<u>Def:</u> difference in magnitudes m from ratio of brightnesses b:

$$m_2 - m_1 = 2.5 \log \frac{b_1}{b_2}$$

(star that is x100 **brighter** has by 5 **lower** magnitude)

absolute scale for apparent magnitude

historically defined

(Sirius: -1.5, Sun: -26.72

naked eye easy: <0, limit: <4 or 6?)

<u>absolute magnitude</u> is measure of luminosity = magnitude that star would have at 10 pc distance

Sun: + 4.83 (visual magnitude)

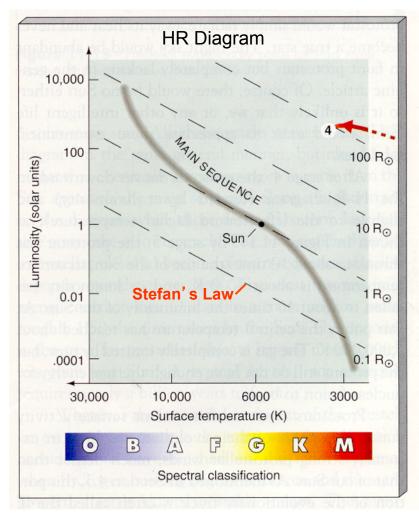
+ 4.75 (bolometric magnitude)

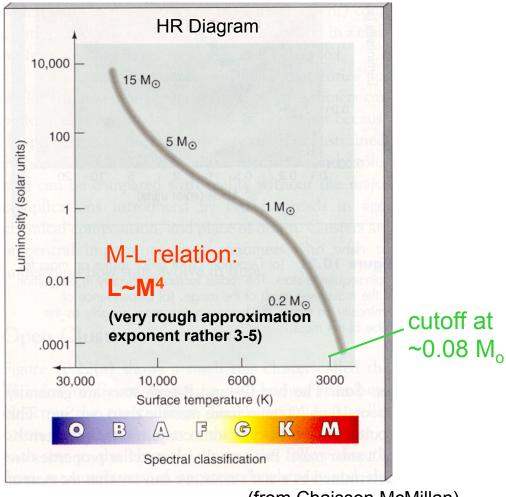
Temperature, Luminosity, Mass relation during core H-burning:

It turns out that as a function of mass there is a rather unique relationship between

- surface temperature (can be measured from contineous spectrum)
- luminosity (can be measured if distance is known)

(recall Stefan's Law L~R² T⁴)





(from Chaisson McMillan)

Mass – Radius relation:

In solar units: $R \sim M^{0.8}$

(10 M sol: 6.3 x R sol, 100 M sol: 40 x R sol) (really exponent is ~0.8 for M<M sol, 0.57 for M>M sol)

Main Sequence evolution:

Main sequence lifetime:

H Fuel reservoir F~M
Luminosity L~M⁴
lifetime
$$\tau_{\rm MS} = \frac{F}{L} \propto M^{-3}$$

Recall from Homework: H-burning lifetime of sun ~ 10¹⁰ years

$$\tau_{\rm MS} \approx \left(\frac{M}{M_{\oplus}}\right)^{-3} 10^{10} \text{ years}$$
 note: very approximate exponent is really between 2 and 3

between 2 and 3

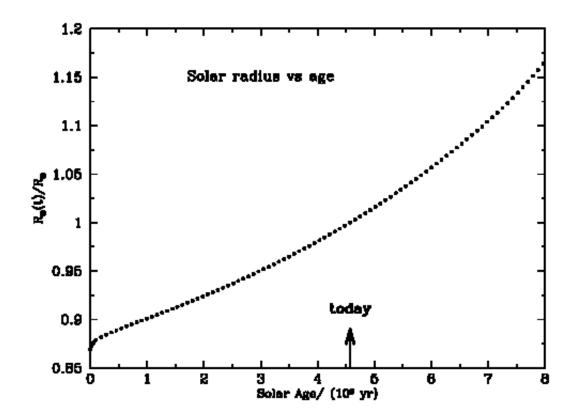
so a 10 solar mass star lives only for 10-100 Mio years a 100 solar mass star only for 10-100 thousand years!

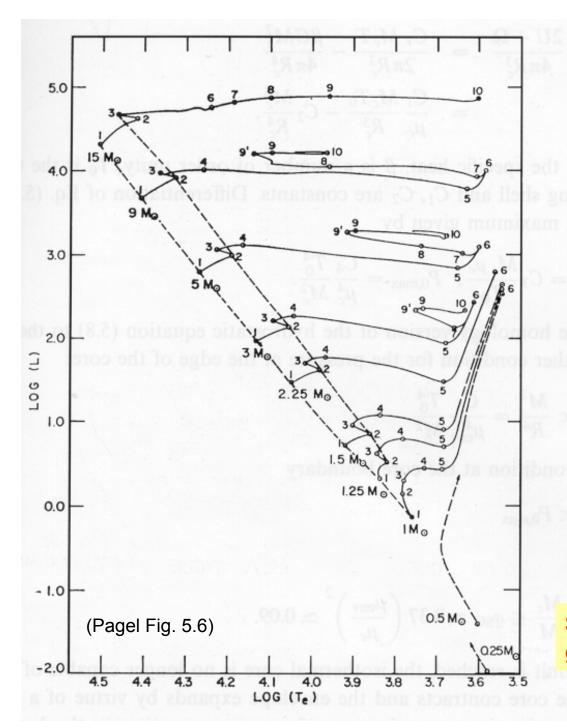
Changes during Main Sequence evolution:

With the growing He abundance in the center of the star slight changes occur (star gets somewhat cooler and bigger) and the stars moves in the HR diagram slightly

main sequence is a band with a certain width

For example, predicted radius change of the sun according to Bahcall et al. ApJ555(2001)990





Zero Age Main Sequence (ZAMS): "1"

End of Main Sequence: "2"

Stellar masses are usually given in ZAMS mass!